



# Can we show that cosmological structures are of quantum mechanical origin?

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# Outline

Based on - J Martin and VV 2015, *arXiv:1510.04038* (PRD) - J Martin and VV 2016, *arXiv:1605.02944* (PRA) - J Martin and VV 2016, *arXiv:1611.01785* (PRA) - J Martin and VV 2017, *arXiv:1706.05001* (PRD)

- Do primordial density fluctuations contain quantum correlations?
- Can we violate Bell inequalities?
- Can we exclude a purely classical explanation?

## **Cosmological Perturbations in Inflation**

Inflation is a high energy phase of accelerated expansion in the early Universe

$$ds^{2} = -dt^{2} + a^{2}(t) d\vec{x}^{2}$$
 with  $\ddot{a} > 0$ 

Quantum vacuum fluctuations are stretched to cosmological scales



**Quantum** Mechanics on **Cosmological** Scales! Can we test it?

Henderson and Vedral 2001, Ollivier and Zurek 2001 Adesso, Bromley, Cianciaruso 2016



Idea: Find two ways to calculate the mutual information between A and B that coincide for classical correlations but may differ in quantum systems

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$$\begin{split} \hat{\Pi}_{j} : \text{complete set of projectors defined on } \mathcal{E}_{B} \\ \hat{\rho} \to \hat{\rho} \hat{\Pi}_{j} / p_{j} \ \text{ with probability } p_{j} = \text{Tr} \left( \hat{\rho} \hat{\Pi}_{j} \right) \text{ and } \rho_{A;\hat{\Pi}_{j}} = \text{Tr}_{B} \left( \hat{\rho} \hat{\Pi}_{j} / p_{j} \right) \\ S(A|B) = \sum_{j} p_{j} S \left( \rho_{A;\hat{\Pi}_{j}} \right) \end{split}$$

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 $\mathcal{I} = S(A) + S(B) - S(A,B)$  $\mathcal{J} = S(A) - S(A|B)$  with respect to measurements  $\hat{\Pi}_j$ 

$$\delta(A,B) = \min_{\{\hat{\Pi}_j\}} \left( \mathcal{I} - \mathcal{J} \right)$$

### **Quantum Discord of Cosmic Inflation**

One scalar degree of freedom:  $v \propto \zeta$  (curvature perturbation)  $\propto \delta T/T$  (CMB T° fluctuation)

• 
$$|\Psi\rangle = \bigotimes_{k \in \mathbb{R}^{3+}} |\Psi_k\rangle$$
 with  $|\Psi_k\rangle = \frac{1}{\cosh r_k} \sum_{n=0}^{\infty} e^{2in\varphi_k} (-1)^n \tanh^n r_k |n_k, n_{-k}\rangle$   
Two-mode squeezed state (Gaussian state)  
•  $\delta(k, -k) = \cosh^2 r_k \log_2 (\cosh^2 r_k)$   
 $-\sinh^2 r_k \log_2 (\sinh^2 r_k)$   
 $\sim 150$  at the end of inflation  
Large degree of quantum correlations!  
Can we design a Bell-type experiment?  
August 2017  $COSMO 17 (Paris)$ 

#### Bell Experiments in the CMB

Spin operators for continuous variables (Larsson 2004)



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Spin operators for continuous variables (Larsson 2004)

$$\begin{split} \hat{S}_{z}(\ell) &= \sum_{n=-\infty}^{\infty} (-1)^{n} \int_{n\ell}^{(n+1)\ell} \mathrm{d}Q |Q\rangle \langle Q| \\ \hat{S}_{x}(\ell) &= \hat{S}_{+}(\ell) + \hat{S}_{+}^{\dagger}(\ell) \\ \hat{S}_{y}(\ell) &= -i \left[ \hat{S}_{+}(\ell) - \hat{S}_{+}^{\dagger}(\ell) \right] \end{split} \text{ with } \hat{S}_{+}(\ell) = \sum_{n=-\infty}^{\infty} (-1)^{n} \int_{2n\ell}^{(2n+1)\ell} \mathrm{d}Q |Q\rangle \langle Q + \ell| \end{split}$$

How do we measure the spin operators?

$$\delta T/T o \zeta_k o v_k$$
 , find  $n$  such that  $n\ell \leq v_k < (n+1)\ell$  and  $S_z = (-1)^n$ 

position measurement only

$$\hat{S}_x$$
 and  $\hat{S}_y$  require to measure the conjugated momentum  $\zeta'_k \sim e^{-r_k} \zeta_k$ , which may be concealed from us for ever...  
See also Campo and Parentani 2005  
Maldacena 2014

#### Can we detect quantum correlations using position measurements only?

COSMO 17 (Paris)

#### Legget-Garg Experiments in the CMB

• Two-time correlators  $C_{ab} = \left\langle \hat{S}_z(t_a, \ell) \hat{S}_z(t_b, \ell) \right\rangle$ 

• Legget-Garg three-strings:  $K_3 = C_{ab} + C_{bc} - C_{ac}$ ,  $K'_3 = -C_{ab} - C_{bc} - C_{ac}$ 

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 $\bigcirc$  But requires to measure  $\zeta$  at three different times...

Can we detect quantum correlations using single-time, position measurements only?

# Back to Discord: Classical States

- Classical states = non-discordant states ( $\delta$ =0) Can we exclude non-discordant states?
- Theorem: the only classical Gaussian states are product states (Adesso and Datta 2010; Rahimi-Keshari, Caves, Ralph 2013; Mista, McNulty, Adesso 2014)



# Conclusions

- **Cosmological perturbations** are placed in a two-mode highly **squeezed state** in the very early Universe
- Such a state has a large **quantum discord**, denoting the presence of **large quantum correlations** between particles created with opposite wave momenta
- In principle, **Bell experiments** can therefore be constructed that would prove that CMB anisotropies are of quantum mechanical origin
- In practice, these experiments require to measure exponentially small quantities (∝ decaying mode), at least in the standard setup
- Legget-Garg inequalities evade this issue but require to measure perturbations at different times
- The CMB cannot have been placed in a **classical Gaussian state**. Current constraints on non-Gaussianities may be already sufficient to exclude non-discordant states!