## Small field models with



## P

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## Motivation

# The hunt for Primordial Gravitational Waves 

- Alternatives to large field models.
- Fundamental Physics?

Ben-Dayan, Brustein JCAP 1009, 007 (2010)

## Outline

I. Recent Results

- The model
- Covering the plane of interest
- Finding the most probable member
II. Discrepancy between predictions and exact calculation
III.Summary and outlook
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## 5 Degree polynomial model

$$
V(\phi)=V_{0}\left(1+\sum_{p=1}^{5} a_{p} \phi^{p}\right)
$$

$$
\stackrel{\Downarrow}{V(\phi)=V_{0}}\left(1-\sqrt{\frac{r_{0}}{8}} \phi\right.
$$

## Results

$$
\left.+\frac{\eta_{0}}{2} \phi^{2}+\frac{\alpha_{0}}{3 \sqrt{2 r_{0}}} \phi^{3}+a_{4} \phi^{4}+a_{5} \phi^{5}\right)
$$

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## Results

Why look at these models?

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## Results

## 6 Degree polynomial model



## $n_{s}$

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## Results

## 5 Degree polynomial model



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Results
5 Degree polynomial model

I. Recent Results

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## 5 Degree polynomial model


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## OOPS!... Something's wrong

## Discrepancy between predictions and calculations


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## Summary and outlook

## What to take away from all of this?

- Small field models can produce GW.
- There is much more in a potential than the first derivative.
- It ain't over 'till it's over.
I. Recent Results
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II. Discrepancy between analytics and numerics
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## Summary and outlook

## In the works

- Find a better analytic expression from first principles.
- Complete work on $r=0.05$ small field models.
- Other checks on small field models viability.


## Questions?

## Thank you! <br> 



## Some Theory

Observables - in theory

- Observables in potential derivative language

$$
\begin{array}{rlrl}
n_{s} \simeq 1 & -6 \varepsilon_{V, 0}+2 \eta_{V, 0} & \varepsilon=\frac{1}{2}\left(\frac{V^{\prime}}{V}\right)^{2} \\
& +2\left[\frac{\eta_{V, 0}^{2}}{3}-\left(\frac{5}{3}-12 b\right) \varepsilon_{V, 0}^{2}\right. & \eta=\frac{V^{\prime \prime}}{V} \\
& \left.-(8 b+1) \varepsilon_{V, 0} \eta_{V, 0}+\left(b+\frac{1}{3}\right) \xi_{V, 0}^{2}\right], & \xi^{2}=\frac{V^{\prime} V^{\prime \prime \prime}}{V^{2}} \\
n_{r u n} \simeq 16 \varepsilon_{V, 0} \eta_{V, 0}-24 \varepsilon_{V, 0}^{2}-2 \xi_{V, 0}^{2} &
\end{array}
$$



